

Physics 216/116

Lecture V - Instabilities I

- So far: - basic eqns.
- Potential Flow
- low Re flow

Sphere + Potential Flow

General ideas  
Wake

$\gamma$  Stokesian Flow

coming

$\mu$  Blasius Boundary Layer (laminar)

Wakes, drag, lift

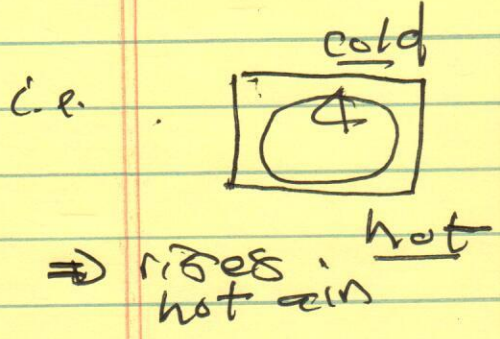
Turbulent Wakes, Turbulence

all: { energy } source for flow is  
body motions (u)

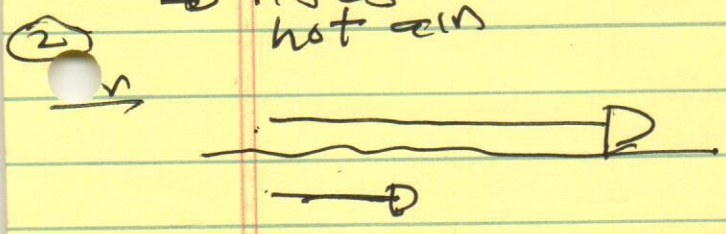
∞∞: Instability → { continuation → KH  
 diversion → RT, RB  
 [stored free energy → fluid motions → chaos, turbulence, dissipation]

⇒ Relaxation - critical stage usually is linear instability

① → stored free energy { reform + small part → growth

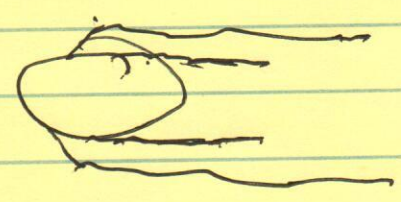


Rayleigh-Bénard convection  
 thermal buoyancy energy  
 →  $DT$  → buoyancy energy

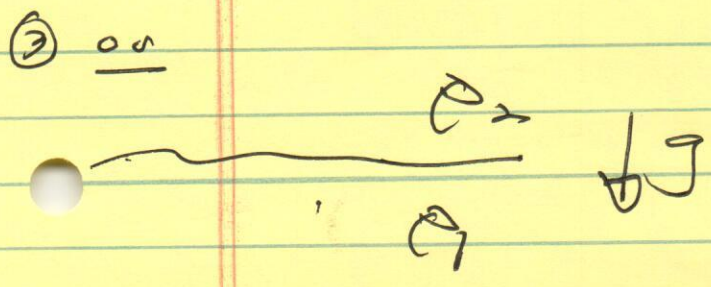


Kelvin-Helmholtz shear flow  
 →  $DV$  → kinetic energy flow shear

relevant to breakdown of wake (after separation)

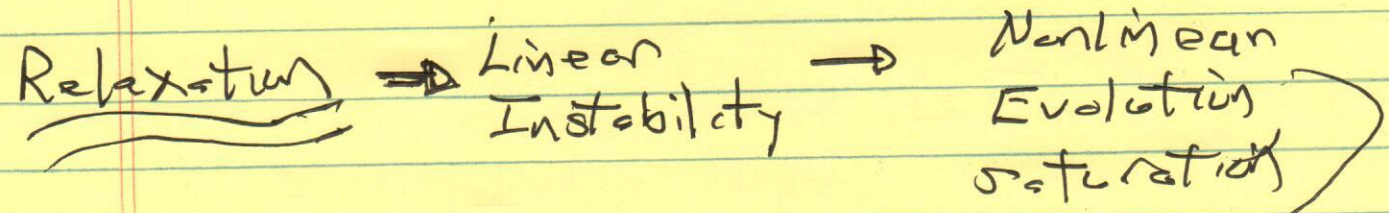


⇒ onset of turbulent wake



Rayleigh-Taylor  
 →  $D\rho + g$  (buoyancy)  
 but, heat not central.  
 → gravitational potential energy.

Real Story/Question :



Final state  
 { cond. dissipation  
 { often turbulent.

Hydro stability  
 is "yugo subject"

c.f. Chandrasekhar

Here: - First step.

[Theory of Hydrodynamic and Hydromagnetic stability]

- 2 classes  $\rightarrow$

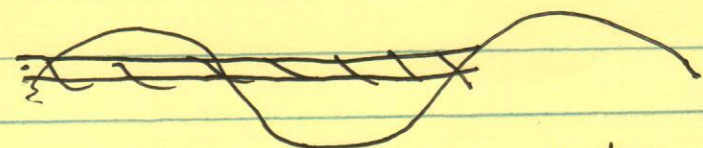
- ① interfacial instabilities  $\rightarrow$  RT, KH
- ② convection  $\rightarrow$  RB

+ homework

# 1.) Interfacial Instabilities

if  $L \ll \lambda$   $\frac{1}{L} = \frac{1}{\rho} \frac{\partial \rho}{\partial z}, \frac{1}{V} \frac{\partial V}{\partial z}$

$k_x L \ll 1$



$\Rightarrow$  treat gradient as held in  $\infty$  interface layer.

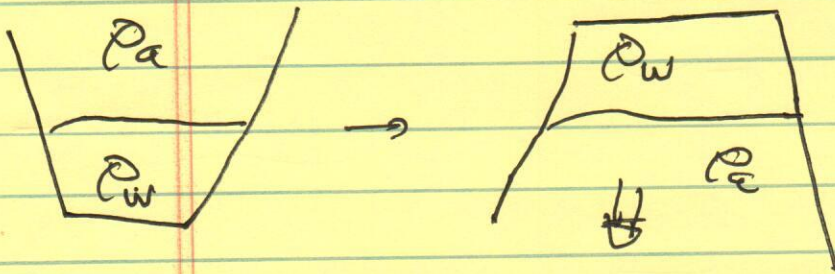
⇒ strategy: - 2 homogeneous media  
 (+)  
 - matching conditions

⇒ significant overlap with theory of { surface phenomena, droplets, etc.

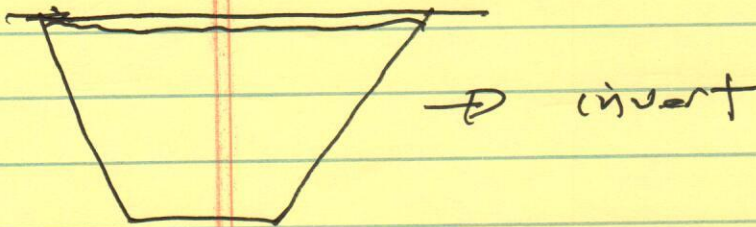
⇒ biophysicist:

n.b. in 'life at low Reynolds number', surface tension relevant.

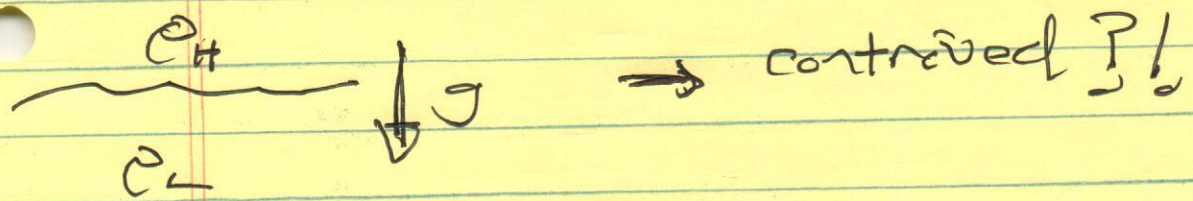
1) Prime Example 1: Rayleigh-Taylor (cf. posted papers, especially Taylor 1950).



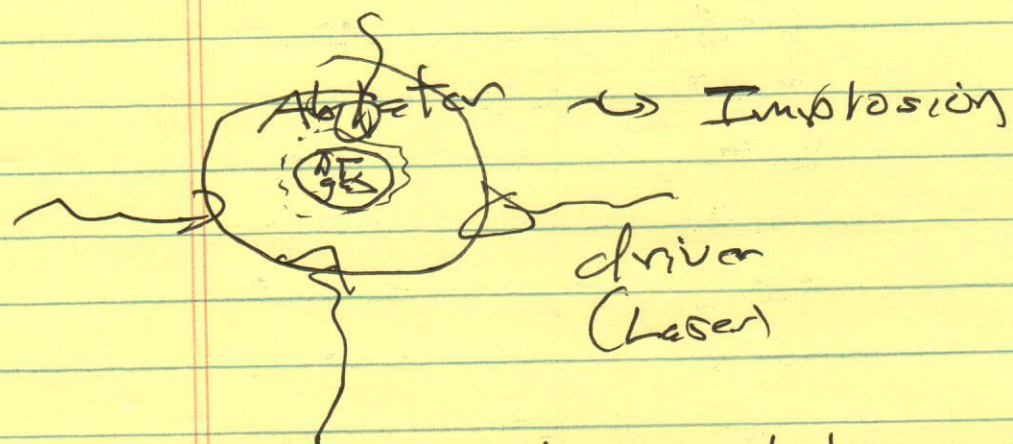
why? ⇒  
 Ripples on surface grow ⇒ R-T. instability



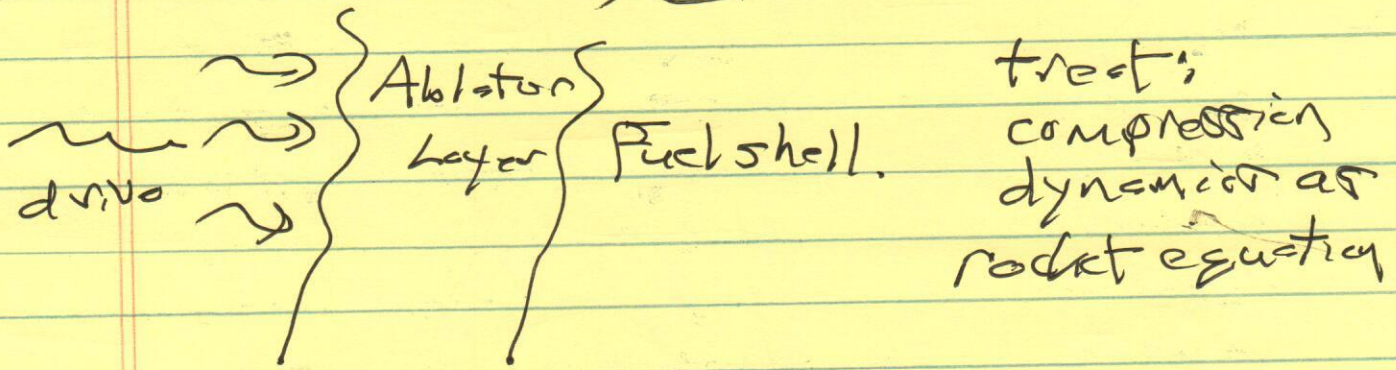
nothing happens!  
 ⇒ cardboard effectively takes  $\gamma$  surf. ⇒ ⊙.



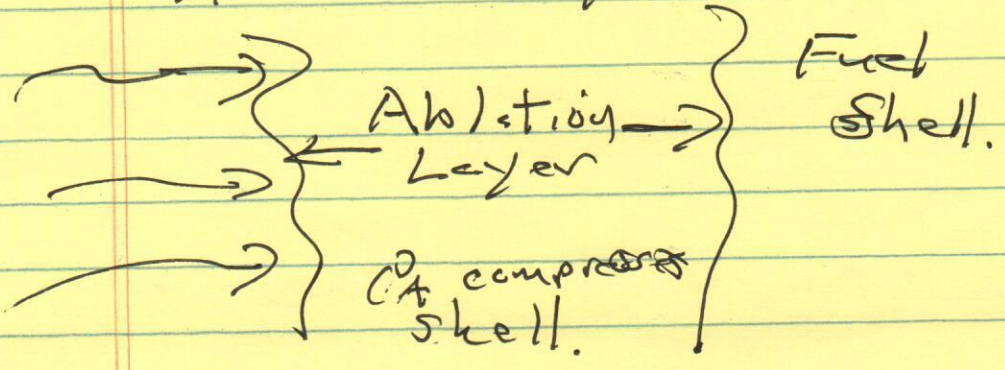
but ICF (controlled and otherwise)



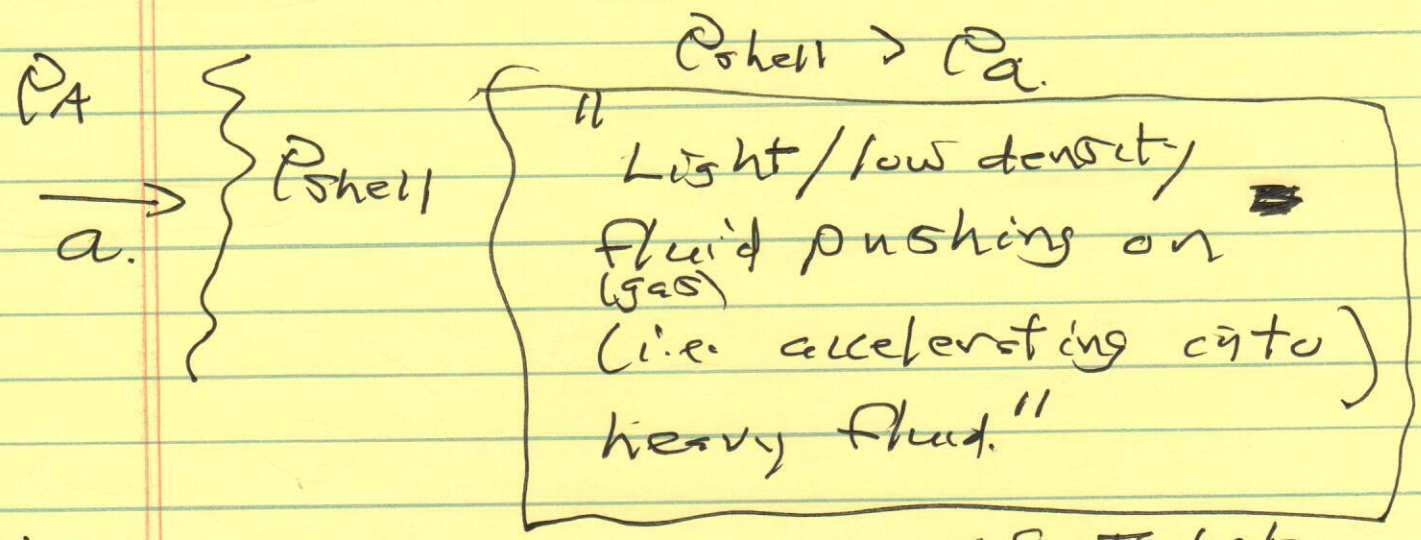
⇒ ablation-driven rocket: driver implosion



⇒ ~~drive~~ drive causes ablation layer to heat and expand, thus compressing inner fuel layer



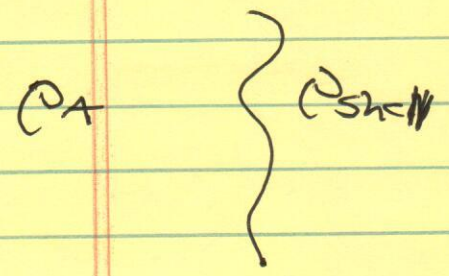
Consider situation:



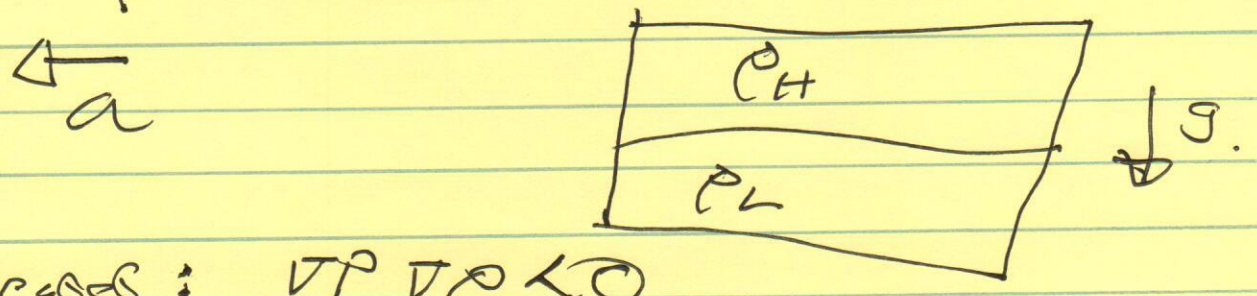
i.e.

in frame of ablator:

c.f. Taylor's paper.

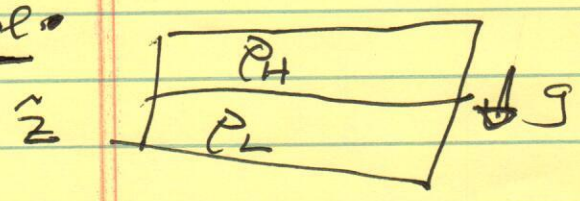


=> equivalent to



Both cases:  $\nabla \rho \nabla \rho < 0$

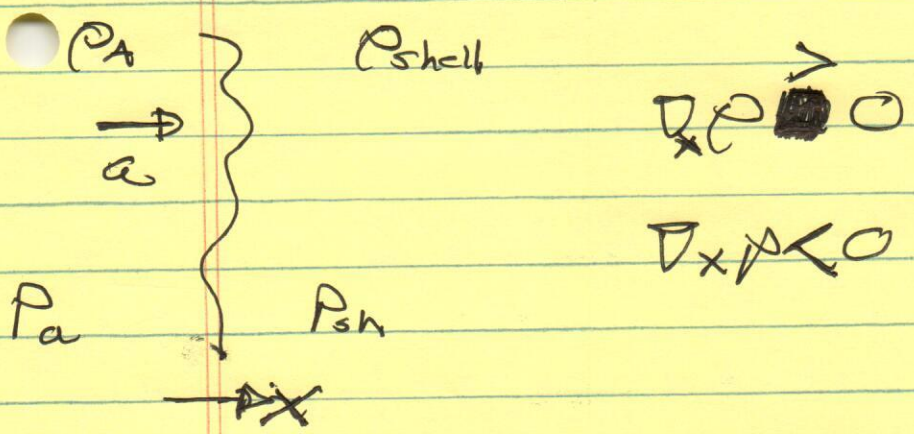
i.e.



$$\nabla \rho > 0$$

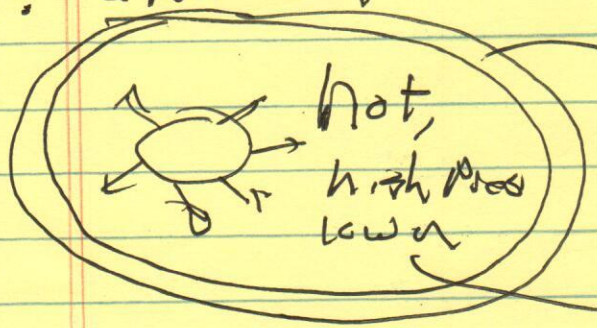
$$\nabla \rho < 0$$

$$\underline{\nabla \rho = -\rho g}$$



$P_a > P_{shell}$       i.e. both  $\nabla \rho < 0$

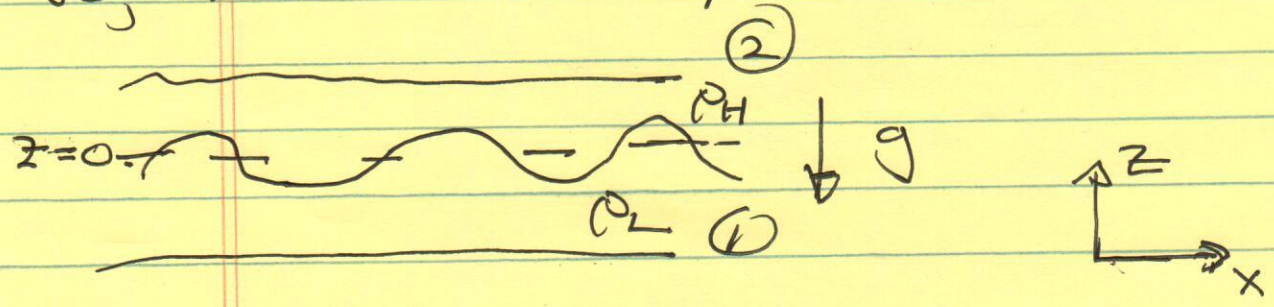
n.b.: also supernovas



→ colder, high density material

→ hot, low density high pressure

so, hereafter: simple case



-  $\nabla \cdot \underline{v} = 0$       i.e. ( $\gamma$  <  $\rho$  etc.)

- ideal fluid (add visc. in HW)

Equilibrium

$$\frac{\partial \underline{v}}{\partial t} + \underline{v} \cdot \nabla \underline{v} = - \frac{\nabla p}{\rho} + \rho \underline{g}$$

$$\nabla^2 p = 0$$

$\rho$  const.  
 $\underline{v} \rightarrow 0$

→ vert,  $\frac{\partial^2 p}{\partial z^2} = 0$

$$p = p_0 + p' z$$

out  $dp/dz = -\rho g$

$$p_2' = -\rho_2 g$$

$$\Delta p > 0$$

$$p_1' = -\rho_1 g$$

$g$  only  $\downarrow$



at interface ( $k_x, k_z \ll k \ll 1$ ), vorticity localized at interface. So treat fluid as irrotational

$$\underline{\omega} = 0, \quad \underline{v} = \underline{\nabla} \phi$$

$$\underline{\nabla} \cdot \underline{v} = 0 \quad \nabla^2 \phi = 0$$

$$\rightarrow \begin{matrix} e^{-kz} \\ \leftarrow \text{---} z=0 \\ e^{+kz} \end{matrix}$$

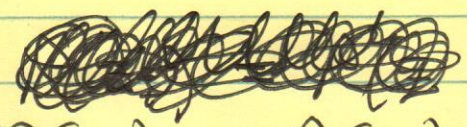
$$\phi = \sum_k e^{ikx} \phi_k(z)$$

$$(\nabla^2 / \partial z^2 - k^2) \phi_k(z) = 0$$

$$\phi_k = \begin{cases} e^{-kz} & z > 0 \\ e^{+kz} & z < 0 \end{cases} \quad (k > 0)$$

at interface ( $z=0$ ) = matching conditions

→ Pressure balance across interface



$$P(\sigma_+) = P(\sigma_-)$$

(else interface in motion on acoustic time scales)

~~V<sub>z</sub>(0<sub>-</sub>) = V<sub>z</sub>(0<sub>+</sub>)~~ V<sub>z</sub>(0<sub>+</sub>) = V<sub>z</sub>(a)

$$\left. \frac{\partial \phi}{\partial z} \right|_2 = \left. \frac{\partial \phi}{\partial z} \right|_1$$

$z \rightarrow 0$                        $z \rightarrow 0$

d.e.

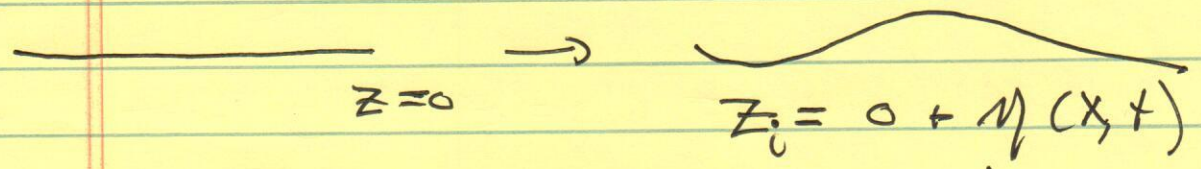
$$\int_{0_-}^{0_+} \left[ \frac{\partial^2 \phi}{\partial z^2} - k^2 \phi \right] dz = 0$$

Note: V<sub>z</sub> b.c. in med & extly forces

$$-k\phi_2 = k\phi_1 \Rightarrow \phi_2 = -\phi_1$$

What of dynamics?

- interface ripples



- displacement of interface
- $\eta$  specifies interface position

Note:  $\phi = \phi(x, z, t)$   
 $= \phi(x, 0+t, t)$   
 $\approx \phi(x, 0, t)$

{ at interface position  
 linear theory

de. linear theory  $\left\{ \begin{array}{l} \phi(x, z, t) \rightarrow \phi(x, 0, t) \\ k\eta \ll 1 \end{array} \right.$

Now must account for force of gravity with displaced interface in Bernoulli's equation:

$$\rho \left( \frac{\partial v}{\partial t} + v \cdot \nabla v \right) = -\nabla p - \rho g \underline{\underline{z}}$$

$$\rho \left( \frac{\partial v}{\partial t} + \nabla \left( \frac{v^2}{2} \right) - \cancel{v \times \omega} \right) = -\nabla p - \rho g \underline{\underline{z}} \quad (g > 0)$$

$$v = \nabla \phi, \quad v_z = \partial_z \phi$$

$$\int_0^\eta dz v_z = \phi_1, \quad \int_\eta^0 dz v_z = \phi_2$$

$$\frac{\partial \phi}{\partial t} + \frac{(\nabla \phi)^2}{2} = -\frac{p}{\rho} - g \eta$$

absent gravity,  $\rho = -\rho \partial\phi/\partial t$

$$\phi = -gz - \frac{\partial\phi}{\partial t}$$

and ~ at surface.

$$\rho = -\rho\eta - \rho \frac{\partial\phi}{\partial t}$$

and finally have equation/dynamic boundary condition for displacement:

$$\frac{d\eta}{dt} = \left. \frac{\partial\phi}{\partial z} \right|_0 = v_z$$

⇒

$$\frac{\partial\eta}{\partial t} + \underbrace{v \cdot \nabla}_{\nabla\phi} \eta = \left. \frac{\partial\phi}{\partial z} \right|_0$$

For stability: linearize

$$\frac{\partial\tilde{\phi}}{\partial t} = -\frac{\tilde{\rho}}{\rho} - g\eta$$

$$\frac{\partial\tilde{\eta}}{\partial t} = \left. \frac{\partial\tilde{\phi}}{\partial z} \right|_0$$

$\rho_1$  }  $\rho_2$   
 $\rho_2$  }  $\rho_1$

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110 noting  $\rho_2 \neq \rho_1$

$$\tilde{\rho}^{(1)} = \tilde{\rho}^{(2)}$$

$$\tilde{\varphi}^{(1)} = -\tilde{\varphi}^{(2)}$$

$$\rho_2 \frac{\partial \tilde{\varphi}^{(2)}}{\partial t} + g \rho_2 \tilde{\eta} = \rho_1 \frac{\partial \tilde{\varphi}^{(1)}}{\partial t} + g \rho_1 \tilde{\eta}$$

$$g(\rho_2 - \rho_1) \tilde{\eta} = \rho_1 \frac{\partial \tilde{\varphi}^{(1)}}{\partial t} - \rho_2 \frac{\partial \tilde{\varphi}^{(2)}}{\partial t}$$

$$= (\rho_1 + \rho_2) \frac{\partial \tilde{\varphi}^{(1)}}{\partial t}$$

$$\left\{ \begin{aligned} \frac{\partial \tilde{\varphi}^{(1)}}{\partial t} &= g \left[ \frac{(\rho_2 - \rho_1)}{(\rho_1 + \rho_2)} \right] \tilde{\eta} \\ \frac{\partial \tilde{\eta}}{\partial t} &= \frac{\partial \tilde{\varphi}^{(1)}}{\partial z} \end{aligned} \right.$$

$$\Rightarrow \left\{ \begin{aligned} \frac{\partial^2 \tilde{\varphi}^{(1)}}{\partial t^2} &= g \left[ \frac{(\rho_2 - \rho_1)}{(\rho_1 + \rho_2)} \right] \frac{\partial \tilde{\varphi}^{(1)}}{\partial z} \end{aligned} \right.$$

using  $\phi \sim e^{-i\omega t} e^{kz} e^{ikx}$

$\Rightarrow -\omega^2 = \left[ g (\rho_2 - \rho_1) / (\rho_1 + \rho_2) \right] k$

$\rho_2 = \rho_H$   
 $\rho_1 = \rho_L$

$\omega^2 = g A k$  ,  $A = \frac{\rho_H - \rho_L}{\rho_H + \rho_L}$  - free energy  
- kinetic  
Atwood #

i.)  $\rho_H = H_2O$   $\lambda \sim 1 \text{ cm}$   
 $\rho_L = \text{air}$   $T_g \sim 1 \text{ sec}$   
(fast!)

ii.)  $\rho_2 = \text{air}$   $\rho_{\text{air}} / \rho_{H_2O} \rightarrow 0$   
 $\rho_1 = \text{water}$

$\omega = \sqrt{k g}$   $\rightarrow$  dispersion relation for surface gravity wave

(stable wave counterpart of R-T)

iii)  $\gamma \sim (gAk)^{1/2}$

→ shorter wavelengths grow faster?

⇒ small scale effects? cut-off, regular?

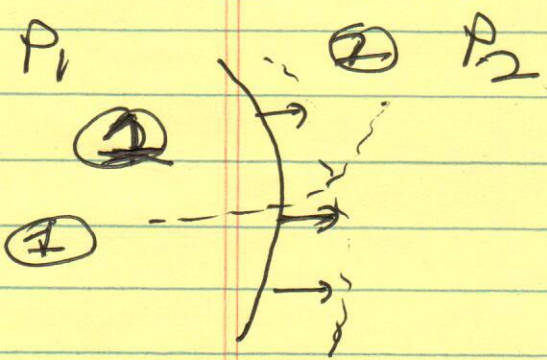
→ viscosity (HW)

→ surface tension

→ finite layer width ( $k_y L_z \geq 1$ )

Surface Tension

→ force due to increase in surface area interface



① expands

↳ isothermal displacement

$$dF = -P_1 dV - P_2 (-dV) + \gamma dA$$

↓ change in free energy      ↓ ① expands into ②      ↓ change in surface area of interface

$$dV = dA \, d\eta$$

↓  
displacement

)  $\eta(x, y)$   
→

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$$dA = \int dx dy \left( 1 + \left( \frac{\partial \eta}{\partial x} \right)^2 + \left( \frac{\partial \eta}{\partial y} \right)^2 \right)^{1/2} - \int dx dy$$

small displacement (slope) :

$$\approx \int dx dy \left( 1 + \frac{1}{2} \left( \frac{\partial \eta}{\partial x} \right)^2 + \frac{1}{2} \left( \frac{\partial \eta}{\partial y} \right)^2 \right) - \int dx dy$$

$$= \int dx dy \frac{1}{2} \left[ \left( \frac{\partial \eta}{\partial x} \right)^2 + \left( \frac{\partial \eta}{\partial y} \right)^2 \right]$$

IBP

$$= \int dx dy \left( - \nabla^2 \eta \right) d\eta$$

↓  
curvature of  
surface displacement

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$$dF = \int \left[ (p_2 - p_1) dA_0 - \nabla^2 \eta dA_0 \right] d\eta$$

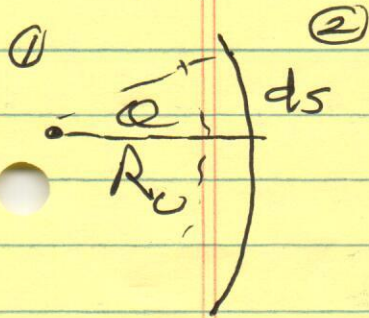


so criteria for equilibrium:

$$P_2 - P_1 = \gamma \sigma^2 \gamma$$

More generally:  $dF = (P_2 - P_1) dA_0 dM + \gamma dA$

Now consider arbitrary (i.e. not "weakly curved" interface):

①   $ds$       ②  $dS = (R_0 + dM) d\theta$

$$= d\theta (1 + dM/R_0) R_0$$

radius of curvature

In general, surface parametrized by 2 radii of curvature,  $R_1, R_2$ :

$$dA = \int dl_1 dl_2 \left(1 + \frac{dM}{R_1}\right) \left(1 + \frac{dM}{R_2}\right) = \int dl_1 dl_2$$

$$= \int dl_1 dl_2 \left(\frac{1}{R_1} + \frac{1}{R_2}\right) dM$$

$$\stackrel{\text{so}}{=} dF = \int \left[ (P_2 - P_1) + \gamma \left(\frac{1}{R_1} + \frac{1}{R_2}\right) \right] dA_0 dM$$

so, for equilibrium with interface (general)

$$\sigma \left( \frac{1}{R_1} + \frac{1}{R_2} \right) = p_1 - p_2 \quad \left[ \text{Laplace's Law} \right]$$

- Given 2-phase equilibrium (separate domains), can use Laplace Law to estimate droplet size for immiscible liquids

c.e.  $p_1 > p_2 \Rightarrow R \sim \sigma / (p_1 - p_2)$

Now, back to R-T, S-W:

$$p_H \gg p_L$$

$$p_H + p_L \rightarrow p_H$$

$$p \rightarrow p - \rho \delta_T \nabla_w^2 \eta$$

$\delta_T \equiv \sigma / \rho$ .  $\rightarrow \sigma$  for each interface  
 c.e. water-air, etc

$\Rightarrow$

$$\gamma_{R-T} = \left( k g A^{\frac{1}{2}} - \gamma_{T \text{ cut-off}} k^{\frac{3}{2}} \right)^{\frac{1}{2}}$$

$$k_{\text{max}} |_{\text{cut-off}} \sim \left( \sigma / \delta_T \right)^{\frac{1}{2}} \rightarrow \text{limits range of unstable modes.}$$

For stable case:

$$\omega^2 = gk + \frac{\sigma}{\rho} k^3$$

} gravity - capillary

↓  
gravity wave  
(long)

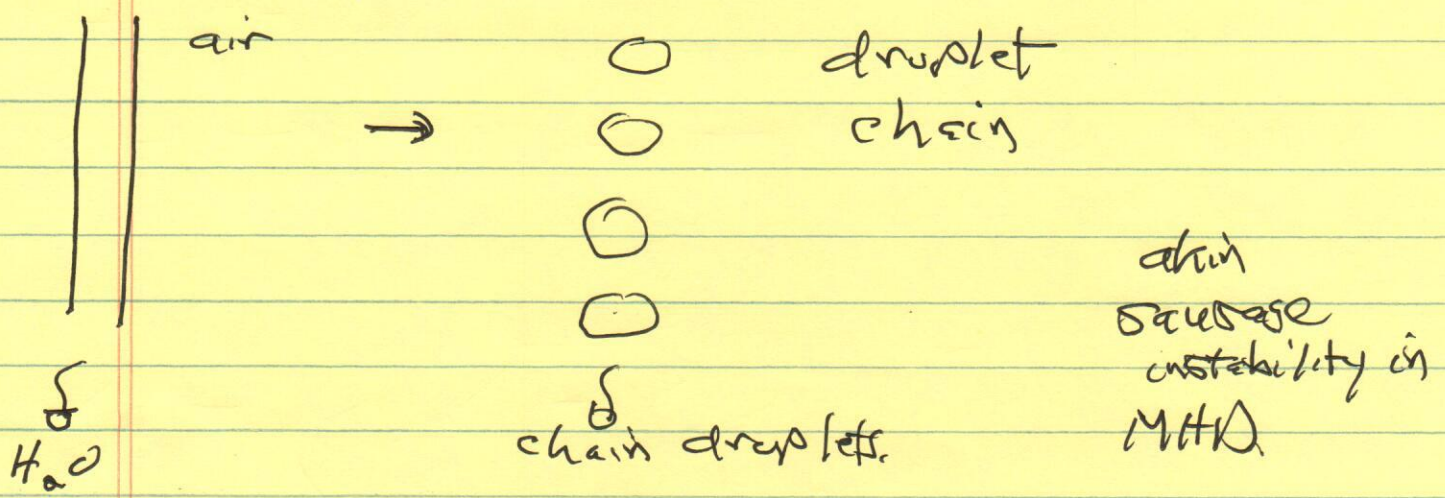
↓  
capillary wave (short)

$$\text{break} \sim \left( \frac{\sigma}{\rho g} \right)^{1/2}$$

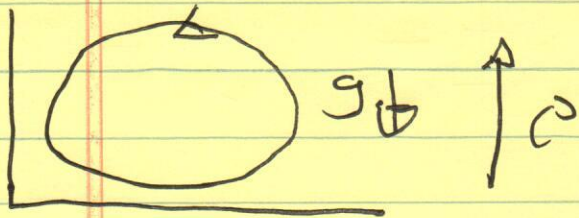
{ in ocean, cross-over at few cm.  
Capillarity important at  $\leq 5$  cm.

N.B.:

Capillarity (S.T.) can induce instability - classic is line of fluid break-up to string of pearls



Note also:



Finite layer thickness

$\Rightarrow$  2D cell - distributed vorticity

$$\omega^2 = \frac{-k_x^2}{k_x^2 + k_z^2} g \frac{1}{L_0} \frac{\partial \psi}{\partial z}$$

$$= -\frac{k_x^2}{k^2} g / L_0$$

$$\gamma^2 = \left( \frac{k_x^2}{k^2} g / L_0 \right)^{1/2} \quad , \quad \text{so } \gamma \uparrow \text{ to flat.}$$

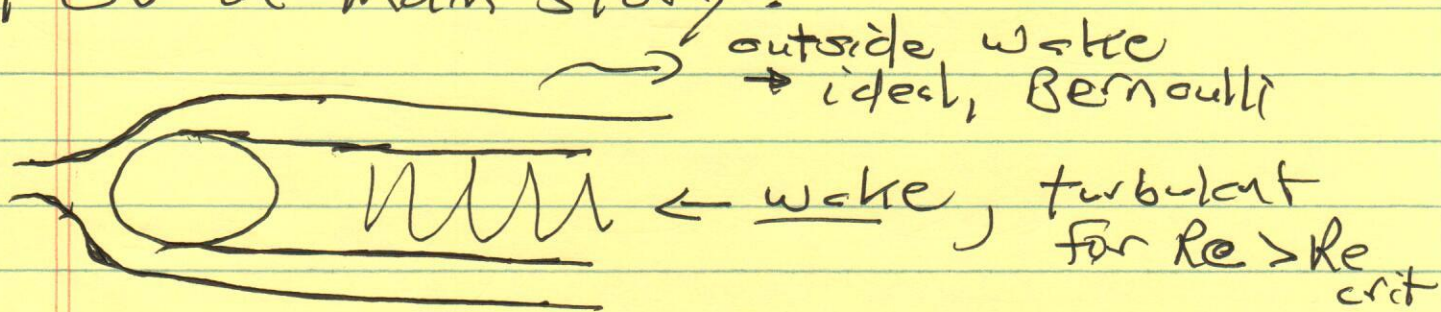
- stable stratification:

$$\omega^2 = \frac{k_x^2}{k^2} g / L_0 \quad \rightarrow \quad \text{internal wave,}$$

$$N^2 \rightarrow \text{BU freq.}$$

→ Kelvin - Helmholtz

Recall OV of main story:



- with ~~no-slip~~ no-slip B.C.'s,  
 $V_n|_{surf} = v_T|_{surf} = 0$

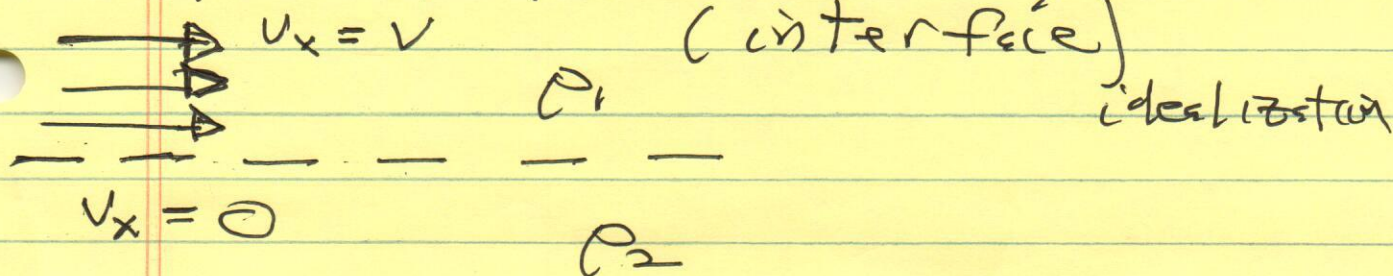
separation happens, wake forms

- separation → instability →  
 turbulence.  
How?

Instability ⇒ Kelvin - Helmholtz

⇒ free energy →  $\Delta V$  — flow shear

⇒ simplification: shear layer  
 (interface)



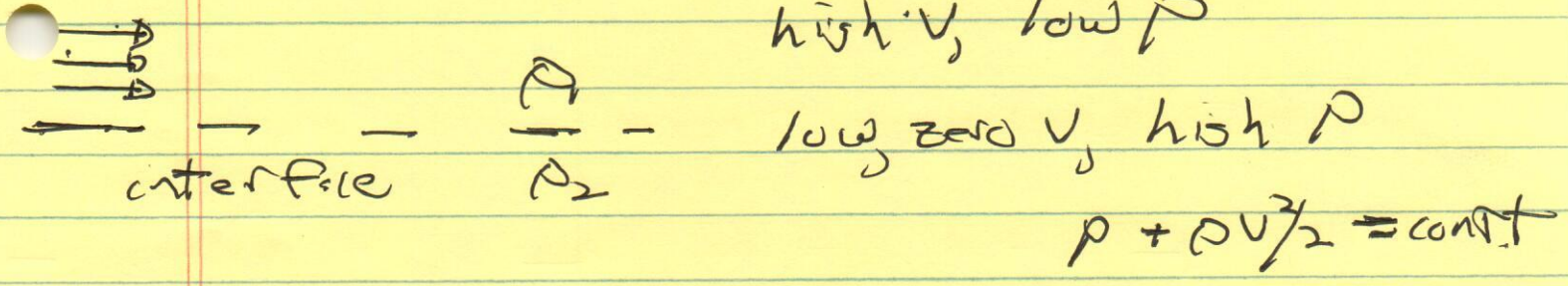
Note:

-  $\nabla V = 0$ , except interface

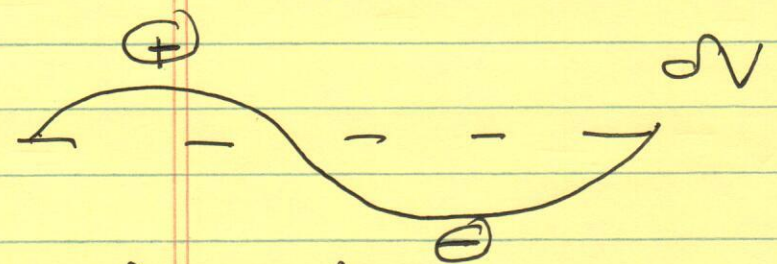
- vorticity  $\partial V_x / \partial z$  localized to interface

$\Rightarrow$  can play game w/  $R-T$ ,  
now with  $V, \eta$ .

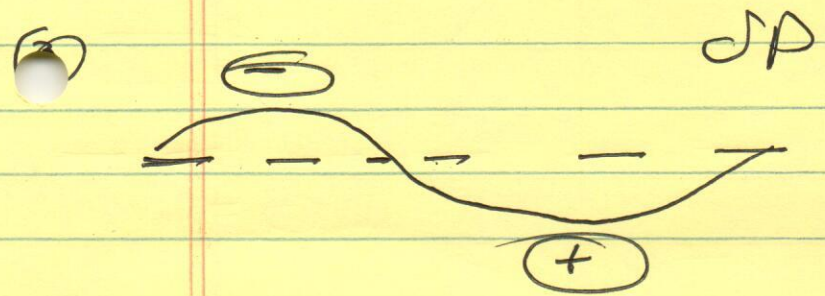
Physical ideas: ①



②  $\delta V$  perturbation  
 $\rightarrow$  ripple interface



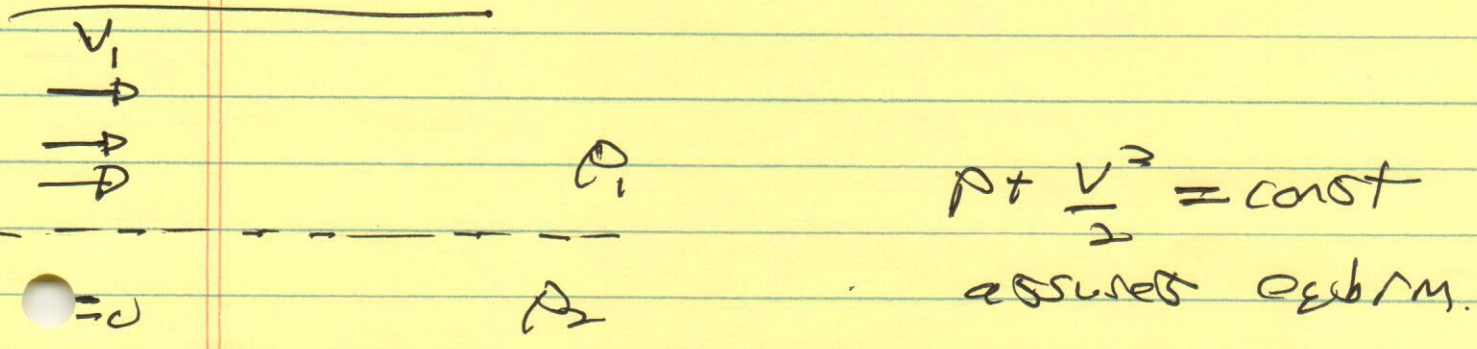
so Bernoulli  $\Rightarrow$



$\delta P < 0 \Rightarrow \delta V > 0$ , further unstable

$\Rightarrow$  KH instability drives viscous mixing via turbulence, mixing, billows, etc.

To calculate:



So, as before:

$$\underline{\nabla} \cdot \underline{v} = 0$$

$$\underline{v} = \underline{\nabla} \phi \quad \underline{\omega} = 0 \text{ except interface}$$

$$\nabla^2 \phi = 0$$

$\Rightarrow$  wave along interface.

$$\phi = \sum_k \phi_k e^{i k x} e^{-k|z|} e^{-i \omega_k t}$$

decays away from interface.

as before;

$$\rightarrow \tilde{P}_1(0_+) = \tilde{P}_1(0_-)$$

$\rightarrow \eta \Rightarrow$  interface ripple/displacement

$$\frac{d\eta}{dt} = v_z|_0$$

and  $v_z(0_+) = v_z(0_-)$

Now,  $\frac{\partial \underline{v}}{\partial t} + \underline{v} \cdot \nabla \underline{v} = -\frac{\nabla p}{\rho}$

$$\rho_2 \left( \frac{\partial \tilde{v}_{z_2}}{\partial t} + v_{z_2} \partial_x \tilde{v}_{z_2} \right) = \left( \frac{-}{+} \right) \nabla_z p_2$$

$$\tilde{v}_{z_2} = \frac{c}{\rho_2 (k v_{z_2} - \omega)} \tilde{P}_2$$

and

A.A.  $\tilde{v}_{z_2}$  so  $c$  from  $\partial_t, \partial_x$  only  
 $\nabla_z p_2$  constant



$$\frac{dM}{dt} = \frac{\partial \tilde{\eta}}{\partial t} + v \frac{\partial \tilde{\eta}}{\partial x} = \tilde{V}_z,$$

$$-c(\omega - kv)\tilde{\eta}_n = \tilde{V}_{zn}$$

||| using Euler/Bernoulli

$$-c(\omega - kv)\tilde{\eta}_n = \frac{-ck\tilde{P}_1}{\rho_1(kv - \omega)}$$

|||

$$\tilde{P}_{1n} = \frac{-\rho_1(kv - \omega)^2}{k} \tilde{\eta}_n$$

$$\tilde{P}_{2n} = \frac{\rho_2 \omega^2}{k} \tilde{\eta}_n$$

△ note sign.  
(opposite signs  $\tilde{P}_1, \tilde{P}_2$ )

and  $\tilde{P}_{1n} = \tilde{P}_{2n} \Rightarrow$

$$-\frac{\rho_1}{k} (kv - \omega)^2 = \frac{\rho_2}{k} \omega^2$$

⇒

$$\omega = kv \left( \frac{\rho_1 + i(\rho_1 \rho_2)^{1/2}}{\rho_1 + \rho_2} \right)$$

⇒

$$\gamma \sim kv \frac{\sqrt{\rho_1 \rho_2}}{\rho_1 + \rho_2} \rightarrow \text{KH growth}$$

note  $\omega_r \sim kv \left( \frac{\rho_1}{\rho_1 + \rho_2} \right)$   
 no "exchange of stabilities" here.

~  $\rho_1 = \rho_2 \quad \gamma = \frac{kv}{2}$

generally  $\gamma \sim k(\Delta v)$ .

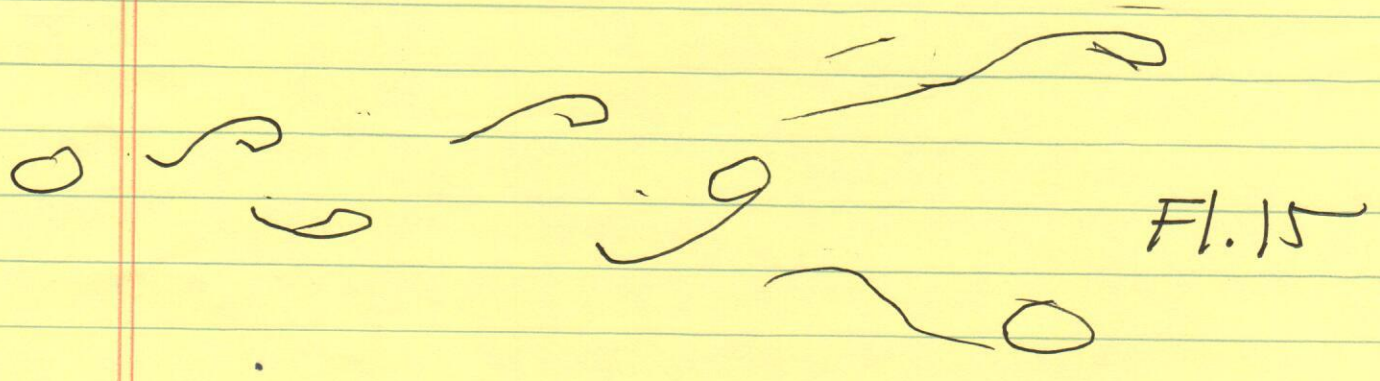
→ what happens?

→ vortex roll-up, billow

F 2.3, F 2.4

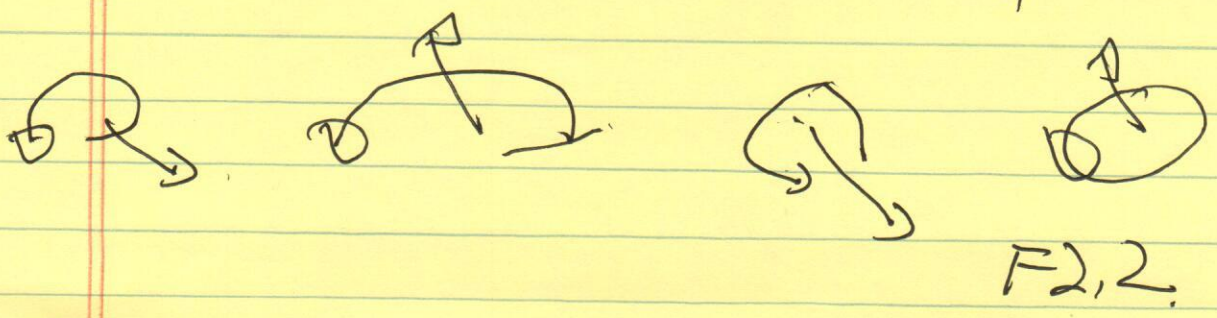


→ Van-Karman vortex street



~~→~~

→ n.b. array vortex lines unstable w/r displacements as shown



Refs:

→ R-T, K-H:

• S. Chandrasekhar, "Hydrodynamic and Hydromagnetic Stability"

→ K-H: Falkovich

→ Surface Wave, Surface Tension (Laplace Law),  
KH: Landau/Lifshitz.

→ see also: G. K. Batchelor "An Introduction to Fluid Mechanics"